

From Science Operations to instrument and Survey Management: Exploring Machine-learning and AI techniques

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Credit: C. Durán/ESO

About

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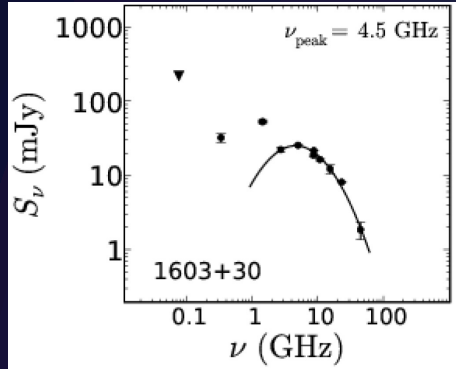


Ph.D. @ ULL
International Research fellow @ INAF
Istituto di Radioastronomia (Bologna)

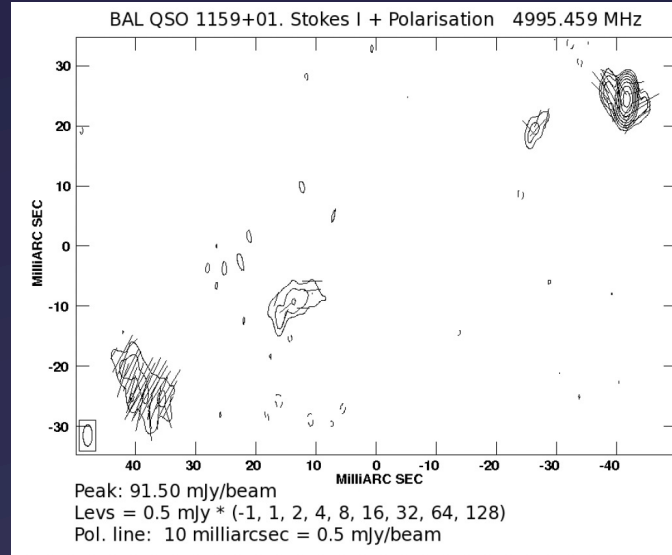


❖ Radio-loud Broad Absorption Line QSOs

Co-directors: K.-H. Mack, J.I. González-Serrano, I. Pérez-Fournón



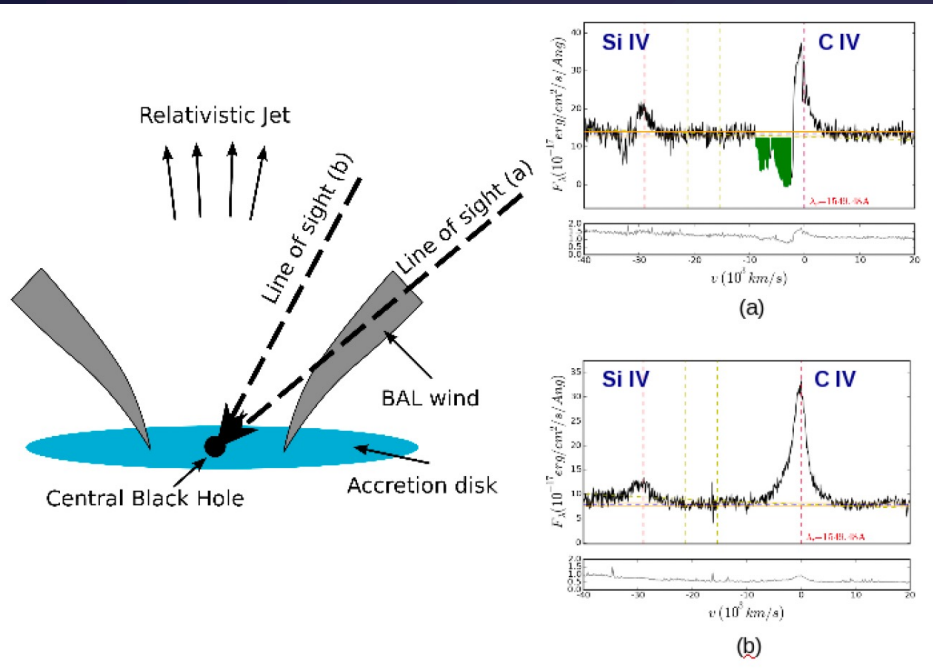
Montenegro-Montes+ 2008,
MNRAS, 388, 1853



BAL QSO team, Bologna

- ❖ Radio properties of BAL QSOs to test the evolutionary vs orientation scenarios
 - ❖ Radio morphology at arcsec and mas scale [compactness, orientation]
 - ❖ Synchrotron age, via SED
 - ❖ Polarisation properties (RM)

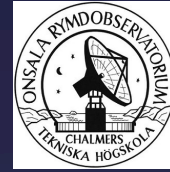
<http://riull.ull.es/xmlui/handle/915/21058>



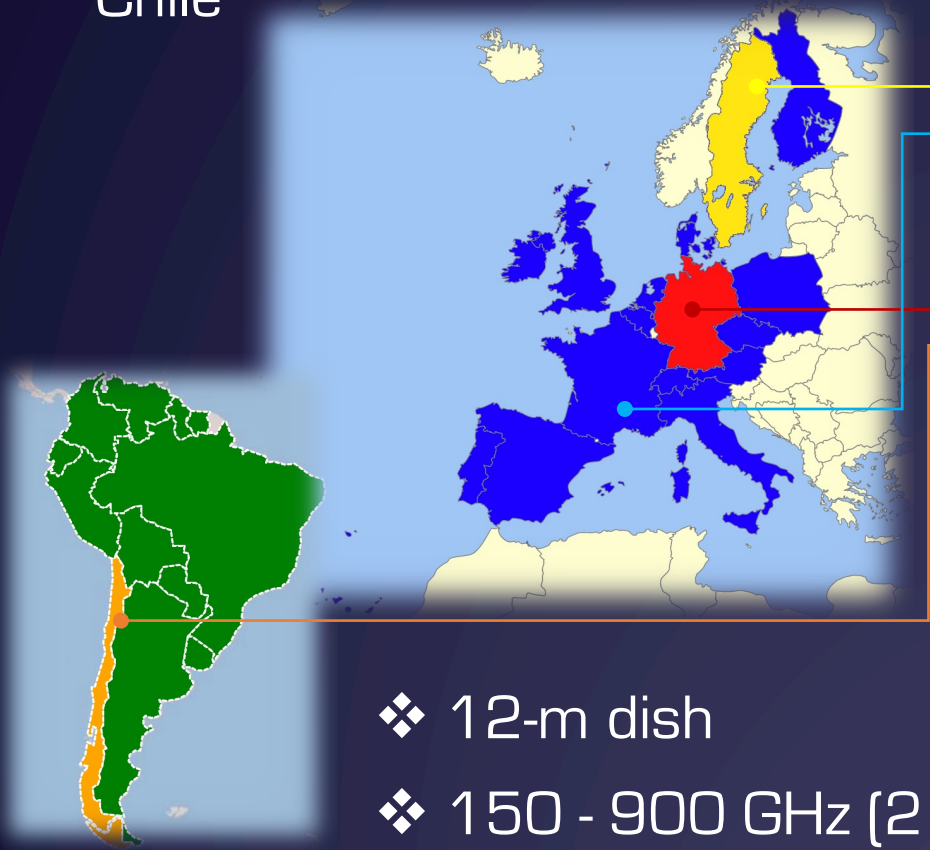
Montenegro-Montes, 2009, PhD T.

❖ European partnership in Chile

Max-Planck-Institut
für
Radioastronomie



APEX



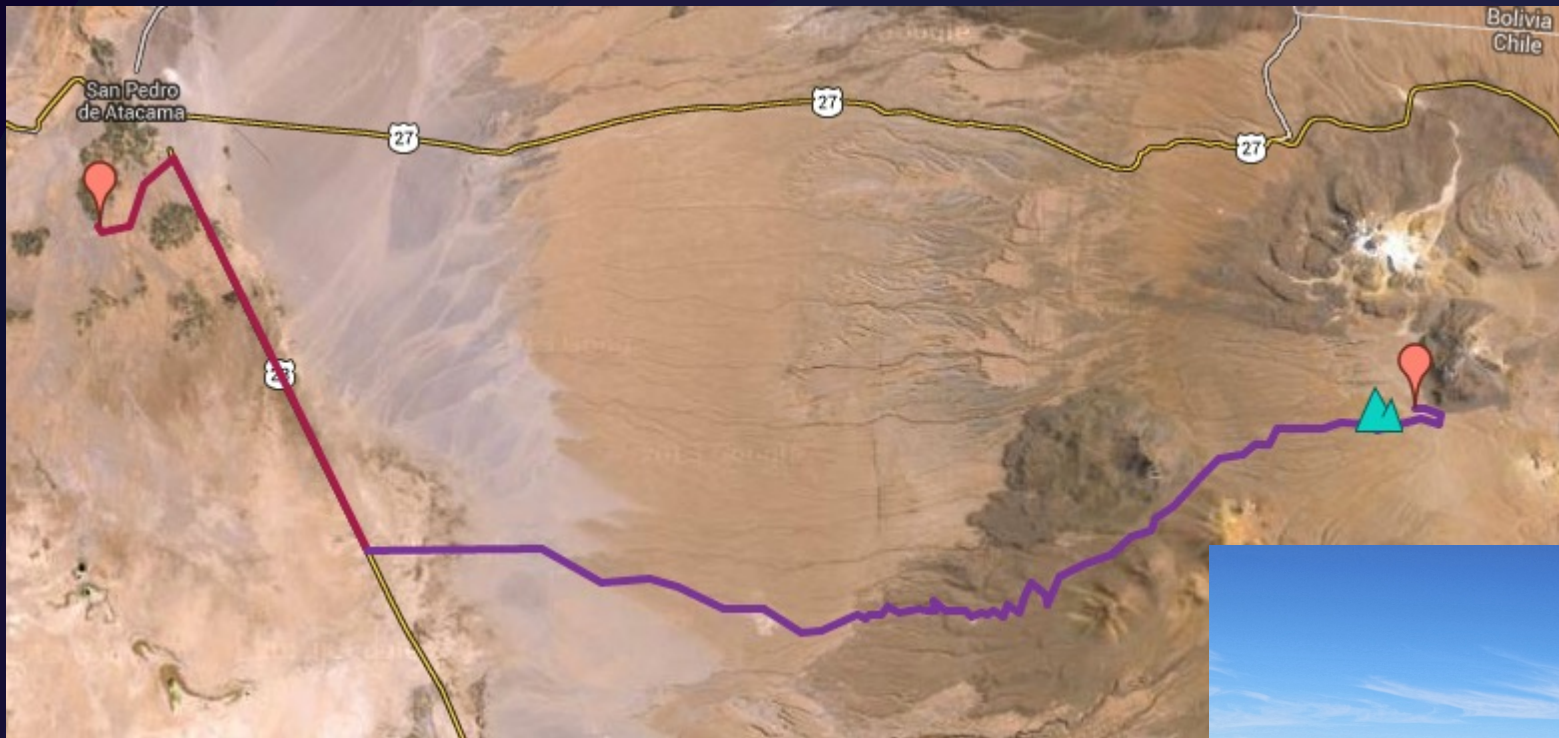
- ❖ 49% MPG and Germany
- ❖ 27% ESO member states
- ❖ 13% Onsala Space Observatory
- ❖ 10% Chile

- ❖ 12-m dish
- ❖ 150 - 900 GHz (2 - 0.3 mm)
- ❖ Operates @5100 m.a.s.l



J. Vieira/APEX (MPIfR/ESO/OSO)

APEX



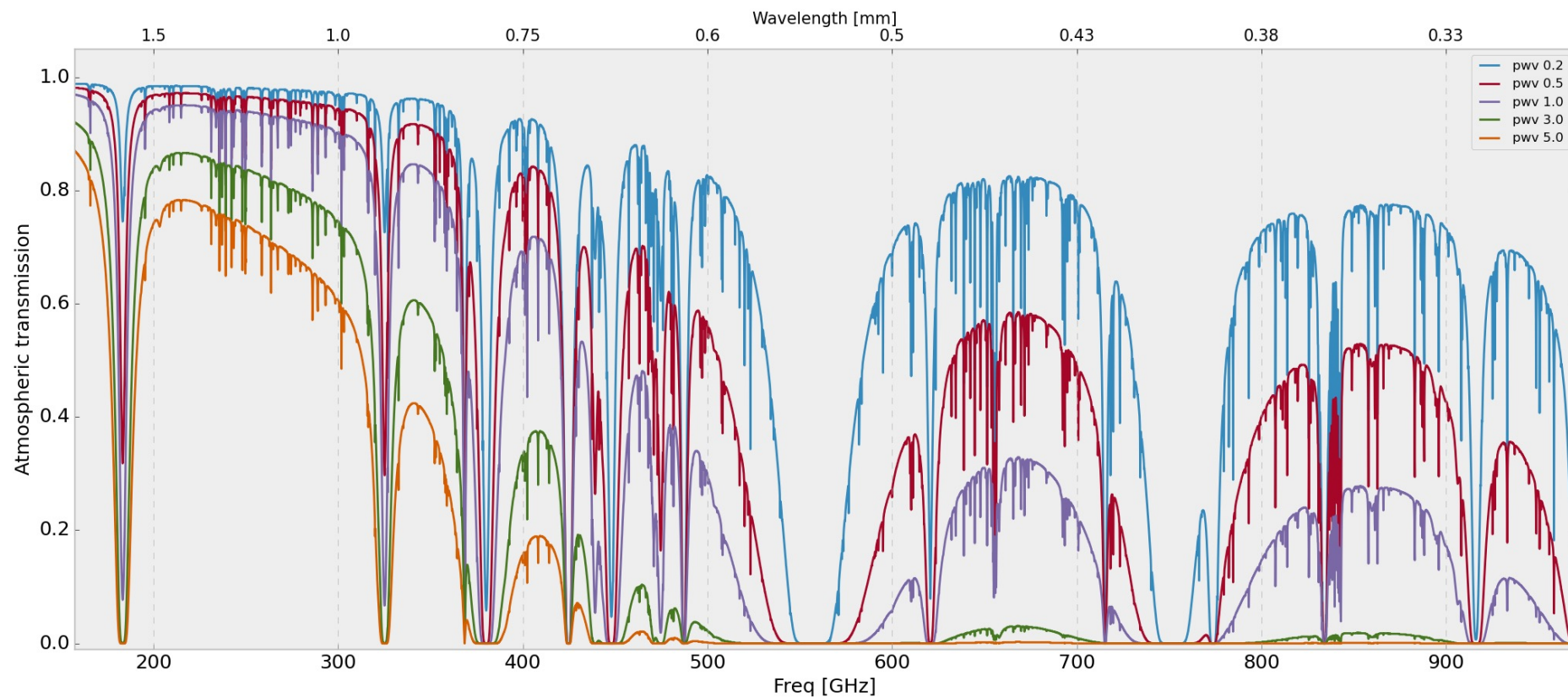
Credit: ESO



IPARCOS Workshop, 16-17 June 2022

Credit: Clem & Adri Bacri-Normier (wingsforscience.com)/ESO

APEX





Credit: ESO/H.H. Heyer

APEX

Location

Llano de Chajnantor. 50 km east of San Pedro de Atacama, Northern Chile

Geodetic Coordinates¹

Latitude: 23° 00' 20".8037 South
Longitude: 67° 45' 32".9035 West
Altitude: 5104.47 m

Geocentric coordinates (ITRF2005)¹

X : 2225039.5297 m
Y: -5441197.6292 m
Z: -2479303.3597 m

Main reflector

Diameter: 12 m
264 aluminum panels with an average panel surface r.m.s. of 5 μ m.
Full surface accuracy: < 15 μ m r.m.s.

Secondary reflector

Diameter: 0.75 m.
Hyperboloidal Aluminum
Polished to an average surface r.m.s. of 2 μ m.

Manufacturer

Vertex Antennentechnik

Mass

125,000 Kg

Mounting

Alt-azimuthal

Primary focal length

4.8 m

Cassegrain focal ratio f/D

8

Beam width

7."8 \times (800 / f [GHz])

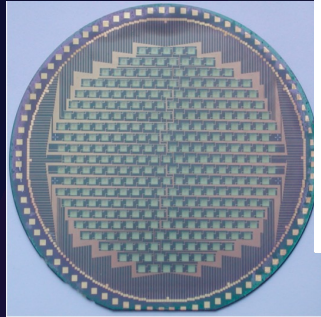
Pointing accuracy (r.m.s.)

< 2" all over the sky
Pointing accuracy on track, 0."6

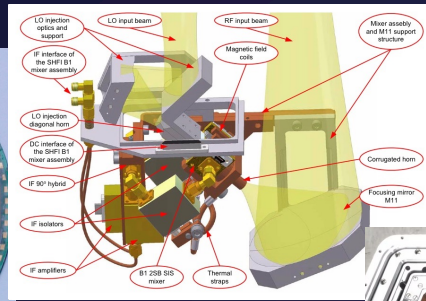
Receiver cabins

2 Nasmyth (A,B) + 1 Cassegrain (C)

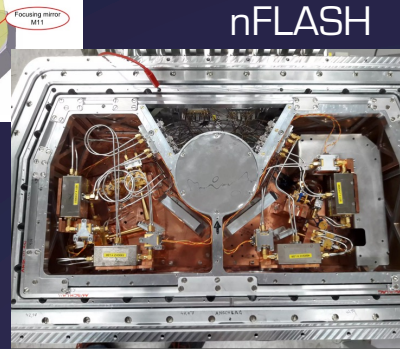
APEX



LABOCA



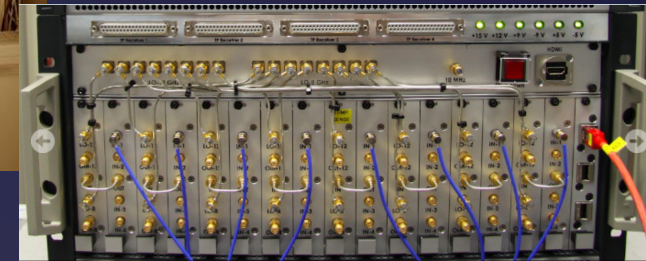
SHeFI



nFLASH



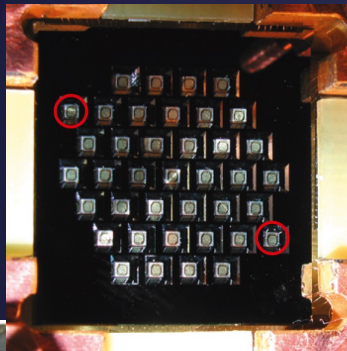
SEPIA



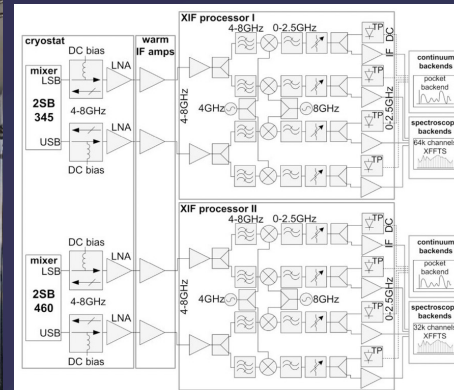
IF/dFFTS4G

Powerful, varied and versatile permanent instrumentation

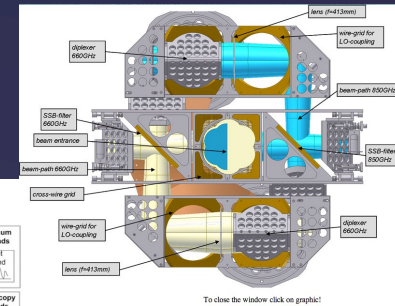
SABOCA



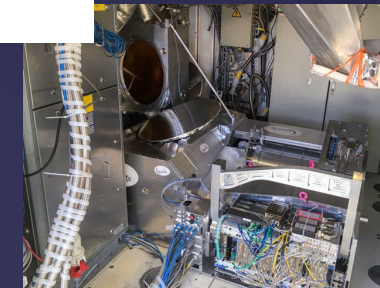
LASMA



FLASH+

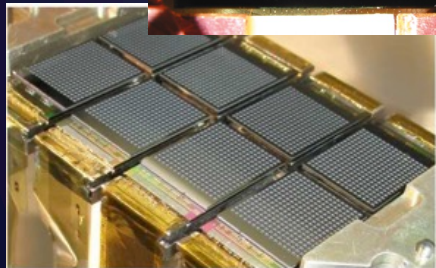


CHAMP+



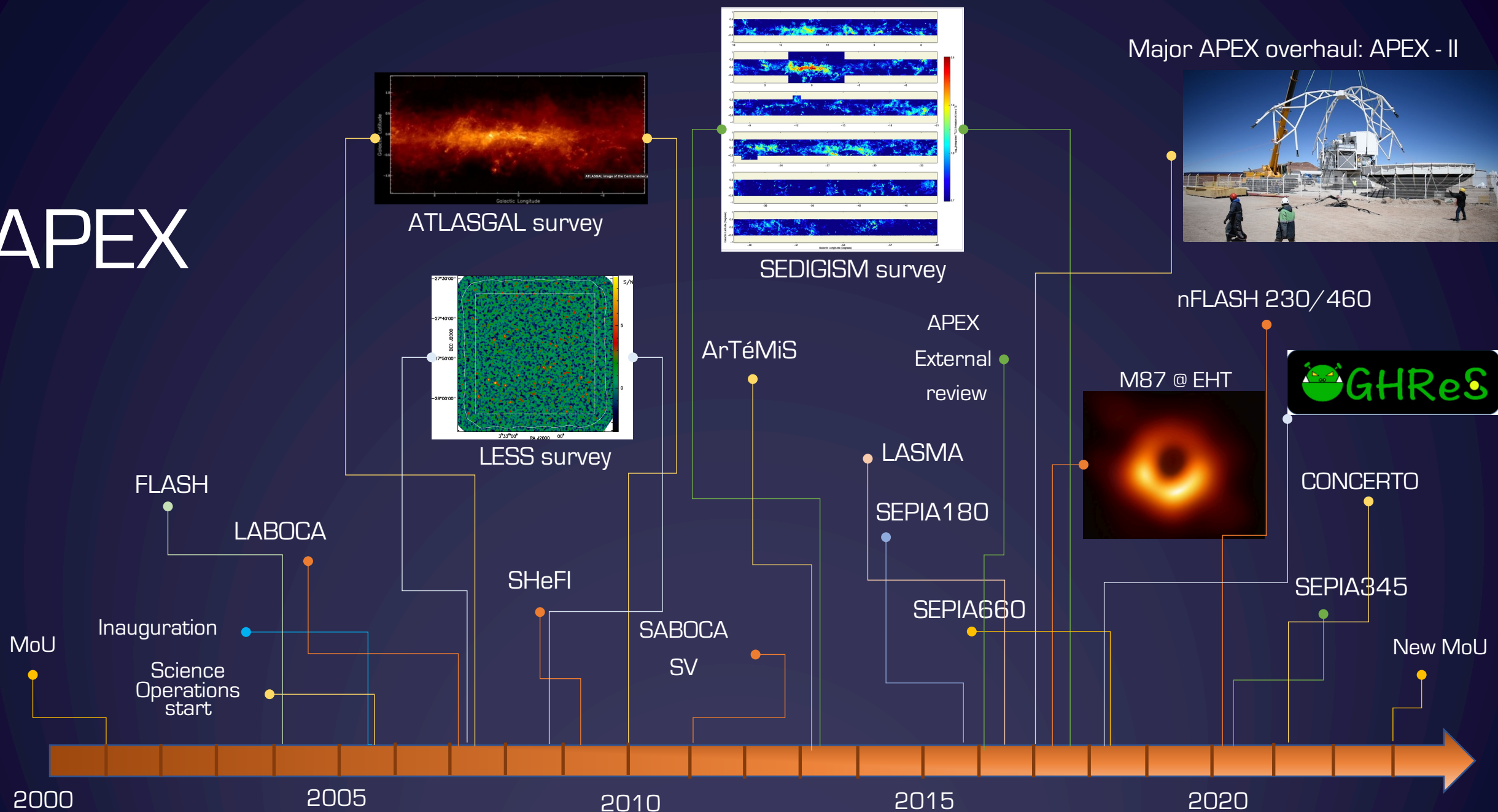
CONCERTO

... and more



ArTéMiS

APEX



Major APEX overhaul: APEX - II

SciOps APEX



- ❖ Night and morning (16h)
- ❖ Sequitor / Chajnantor
- ❖ Oxygenized container

- ❖ Sporadically
- ❖ Chajnantor, 5k
- ❖ No oxygenized container

- ❖ All day
- ❖ Sequitor (night)
- ❖ Chajnantor Morning/afternoon
- ❖ Oxygenized container



- ❖ All day
- ❖ Fully remote (Sequitor)
- ❖ SAS system



SciOps APEX 2018 2022

Principles:

❖ Efficiency

Minimize downtime
Automation

❖ Data Quality

Monitor sub-systems
and data

❖ Service to partners

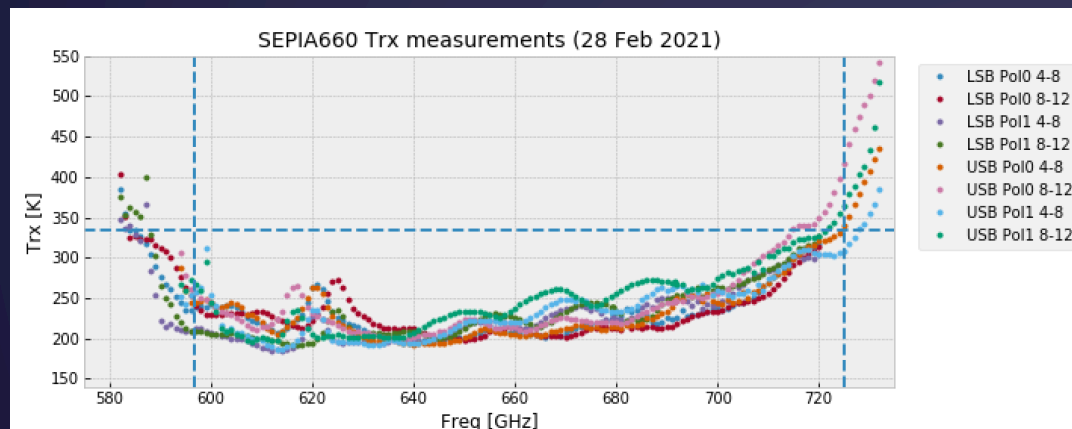
Tools, documentation,
support to instrument
teams

Sciops WP-1000				
Observations WP-1100	Data quality & integrity WP-1200	Antenna and Instruments Performance WP-1300	Stakeholder support WP-1400	Internal development and maintenance WP-1500
Observation planning and setup WP-1110	Data quality assurance WP-1210	Antenna system monitoring WP-1310	Commissioning activities WP-1410	Sciops internal and monitoring tools WP-1510
Antenna and instrument operation WP-1120	Scientific goal assessment WP-1220	Instruments performance monitoring WP-1320	Hardware maintenance activities WP-1420	Internal training WP-1520
Observation logging WP-1130	Scientific data backup WP-1230	Troubleshooting WP-1330	Observing training for visiting astronomers WP-1430	Sciops internal documentation WP-1530
Weather and environment monitoring WP-1140	Science Data transfer WP-1240		Documentation for PIs WP-1440	Group coordination activities WP-1540
			Devel. & maint. of tools for PIs/observers WP-1450	Scientific Work (Santiago) WP-1550

Monitoring and Automation efforts

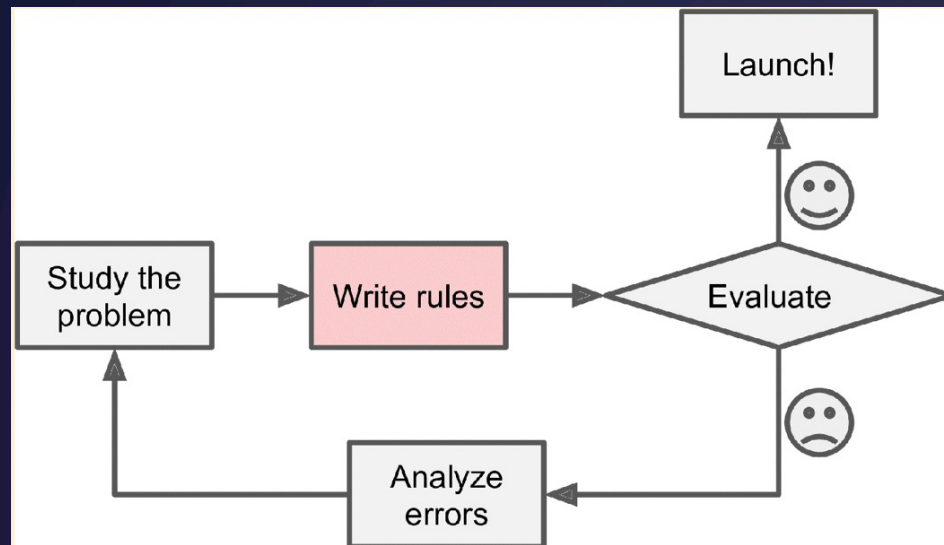
"You cannot manage, what you cannot measure"

- ❖ Monitoring points in as many subsystems as possible/useful (telescope, instruments, aux systems)
- ❖ Data storage on redundant relational databases (MySQL)
- ❖ Graphical dashboards to visualize evolution of KPIs (Grafana)
- ❖ Alarms for deviations from expected values
- ❖ Python scripting to automate observing macros
- ❖ Python/Jupyter to access data and analyse (troubleshooting)
- ❖ Anomaly detection (failures, preventive maintenance)

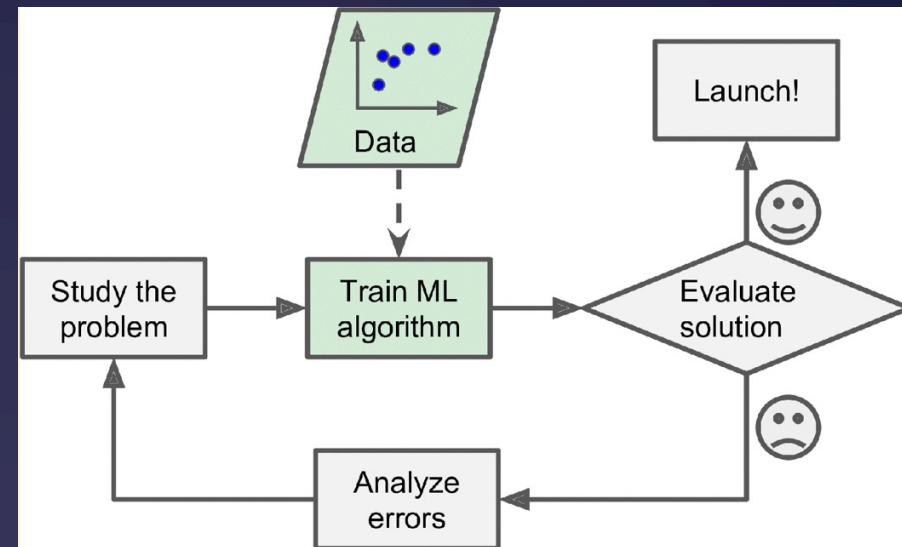


Machine Learning & Artificial Intelligence

❖ Ability to learn from data [without being explicitly programmed].



Traditional approach

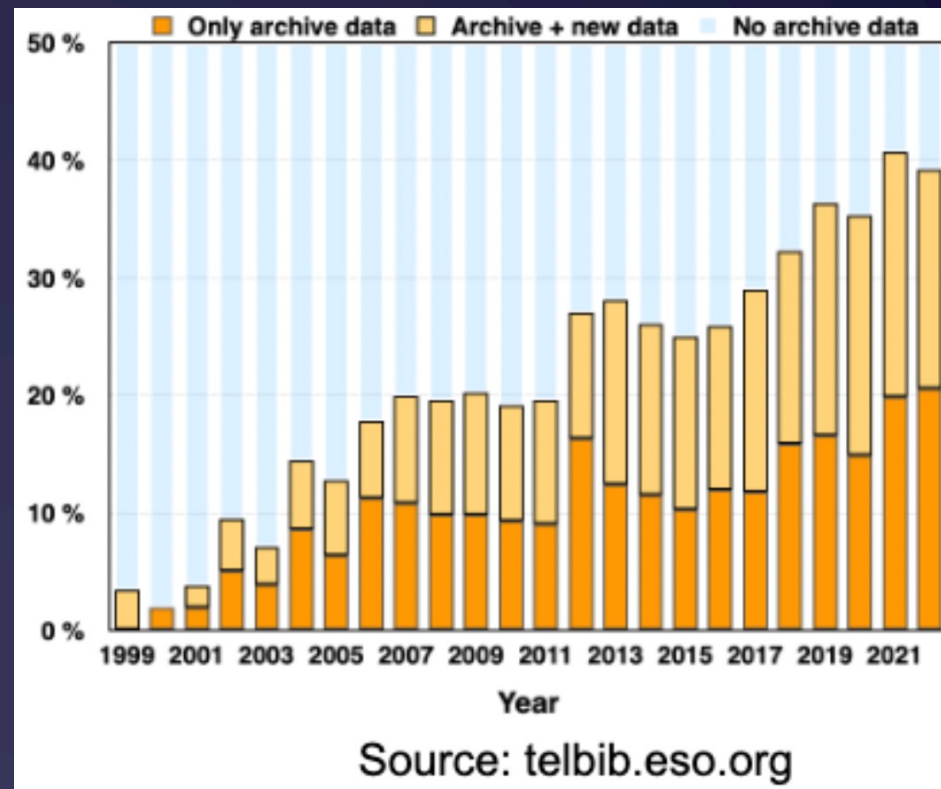


ML approach

Hands-On Machine Learning with Scikit-Learn, Keras, and TensorFlow , 2nd Edition, by Aurélien Géron (O'Reilly). Copyright 2019 Kiwisoft S.A.S., 978-1-492-03264-9."

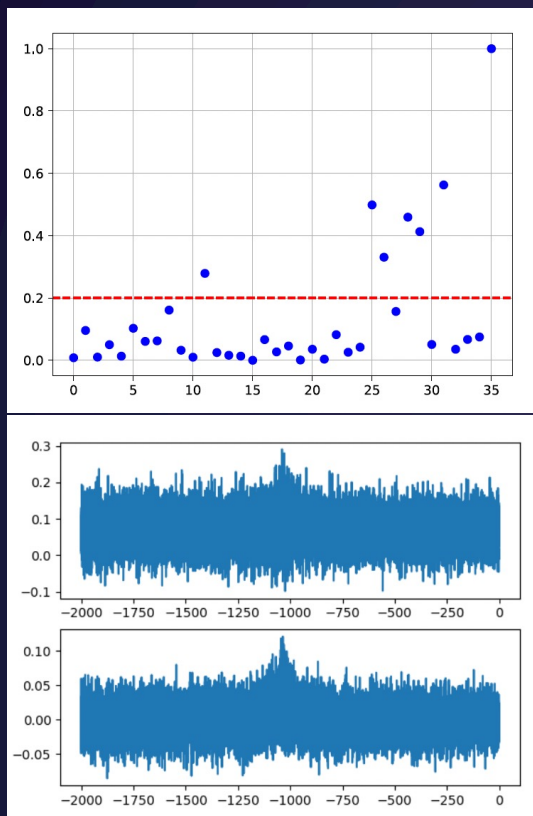
Machine Learning & Artificial Intelligence

- ❖ Successfully used in several scientific fields and contexts, including astronomy
- ❖ Potential for exploitation of public archives and catalogs
- ❖ Era of big data in astronomy (Vera Rubin, former LSST)
- ❖ Ethical problems inherent to the use of AI not so much of an issue for astronomical research



Machine Learning & Artificial Intelligence

❖ Successfully used in several scientific fields and contexts, including astronomy



Anomaly detection in APEX het. subscans
(Local Outlier Factor, `scikit-learn`)

*F. Montenegro, C. Agurto, P. Venegas,
2019. AIA @ ESO*

Science Operations

	Token	weight
seqERR_BAD_QUALITY : "<alias>SENSOR20.last({})...		-11.400035
ccsERR_DB_QUALITY : Point <alias>SENSOR20.last...		-11.383230
... "Maintenance" OBS.START "{}" OBS.NTPL "{}" ...		-8.649286
... Analysis" TPLPRESEQ "GRAVITY_gen_tec_Pola...		-8.649286
WAIT -all (blue)		-8.235087
Setting HWP to: {} encs		-7.677115
Forward(b) SETUP to NGCIR3		-7.555012
Last Reply to 'START' from 'NGCIR3' received: ...		-7.555012
Send command 'START' '{}now' to sub-system 'N...		-7.555012
Cropping NGCIR1 failed!		-7.555012
Enable newdata newdataNGCIR3		-7.555012

Log analysis in Paranal Observatory (NPL + various techniques)

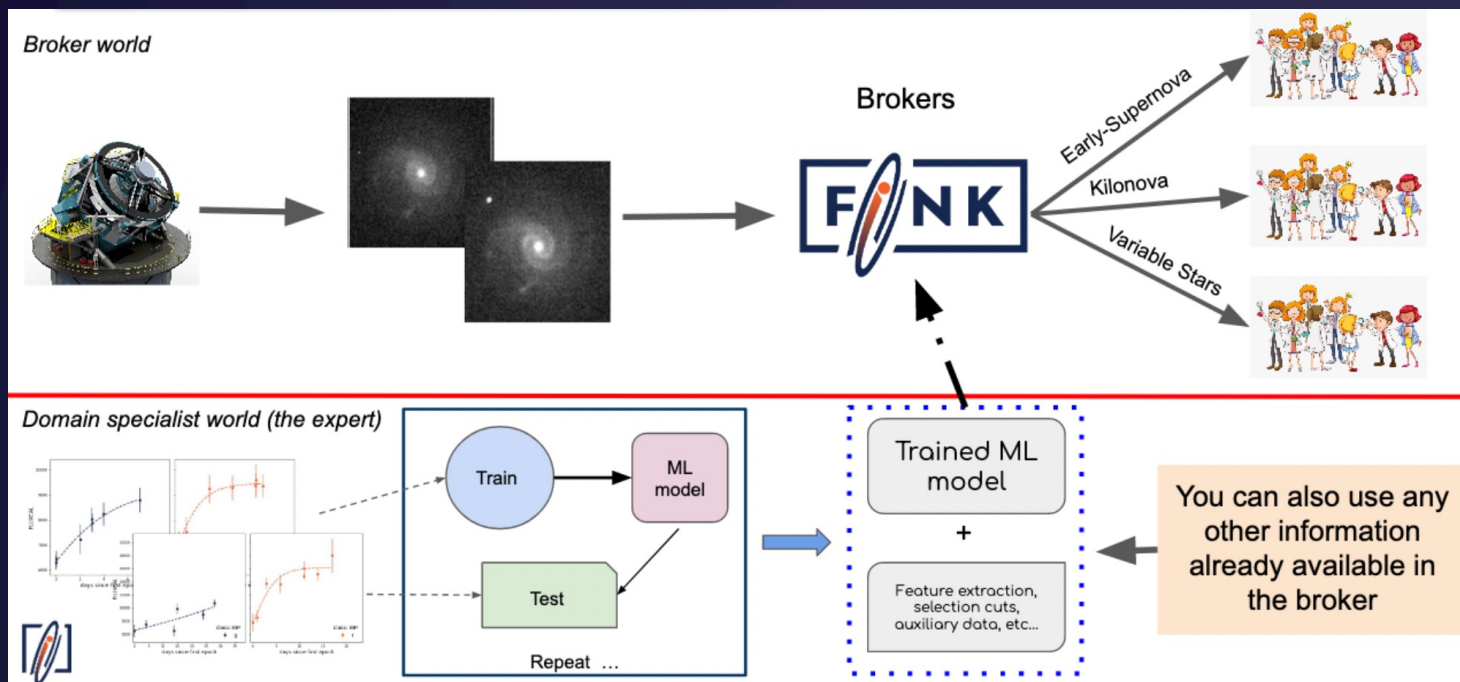
Juan Pablo Gil, 2022 SCIOPS22 Conference

Dataset Generation	Pre-processing	Event Representation	Technique
<ul style="list-style-type: none">- High Level Task<ul style="list-style-type: none">- Observation- Template Exec.- PRESET- SELINS- Split in Traces- Local CSV Files	<ul style="list-style-type: none">- Cleansing: remove bad traces- Tokenization (equivalent to stemming, etc, in NLP)- Vocabulary generation	<ul style="list-style-type: none">- Time series- Tokens- Bag of Words (BoW)- One Hot BoW- Successor Graphs	<ul style="list-style-type: none">- Time series analysis- Naive Bayes- Cluster detection:<ul style="list-style-type: none">- K-Means- SVM- T-SNE- UMAP- Topic Modelling<ul style="list-style-type: none">- LDA- Deep Learning<ul style="list-style-type: none">- Transformers- Sequence Detection

Machine Learning & Artificial Intelligence

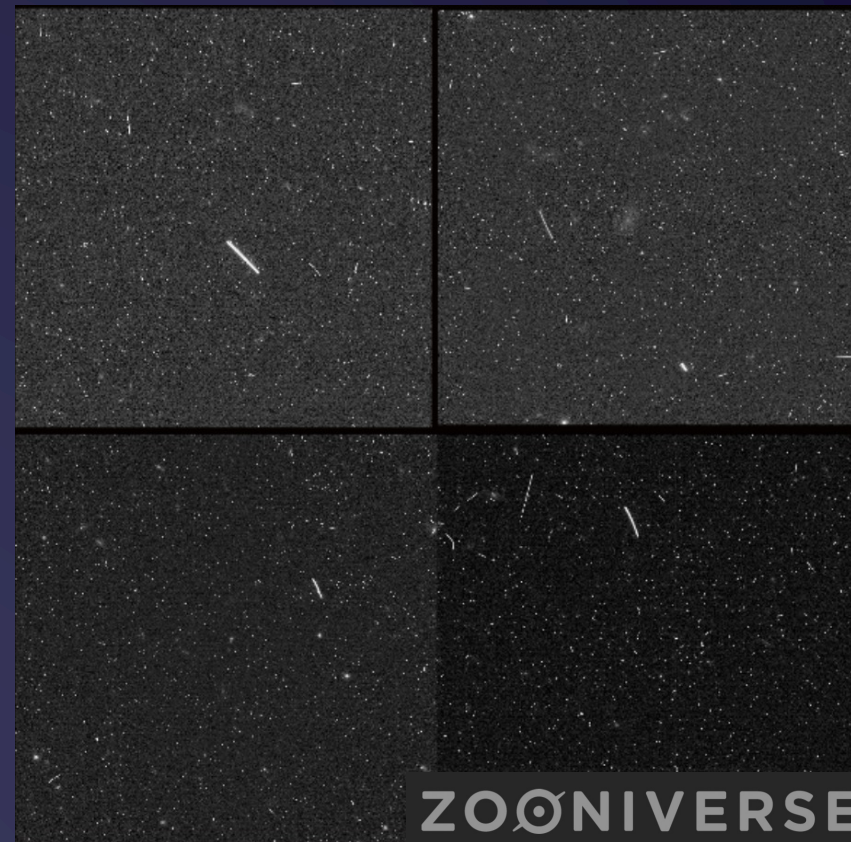
- ❖ Successfully used in several scientific fields and contexts, including astronomy

Classification and Alert systems for transient objects



The Fink broker, E. Ishiida, 2022. DOI 10.5281/zenodo.6554459

Supervised classification of asteroid trails




S. Kruk (ESA).NASA/ESA Hubble Space Telescope

Machine Learning & Artificial Intelligence

❖ Successfully used in several scientific fields and contexts, including astronomy

Spectroscopic observations of the machine-learning selected anomaly catalogue from the AllWISE Sky Survey[★]

 A. Solarz¹, R. Thomas¹, F. M. Montenegro-Montes¹, M. Gromadzki², E. Donoso^{3,4}, M. Koprowski⁵, L. Wyrzykowski², C. G. Diaz^{3,6}, E. Sani¹ and M. Bilicki⁷



Received: 18 May 2020 | Accepted: 14 August 2020

Abstract

We present the results of a programme to search and identify the nature of unusual sources within the All-sky Wide-field Infrared Survey Explorer (WISE) that is based on a machine-learning algorithm for anomaly detection, namely one-class support vector machines (OCSVM). Designed to detect sources deviating from a training set composed of known classes, this algorithm was used to create a model for the expected data based on WISE objects with spectroscopic identifications in the Sloan Digital Sky Survey. Subsequently, it marked as anomalous those sources whose WISE photometry was shown to be inconsistent with this model. We report the results from optical and near-infrared spectroscopy follow-up observations of a subset of 36 bright ($g_{AB} < 19.5$) objects marked as “anomalous” by the OCSVM code to verify its performance. Among the observed objects, we identified three main types of sources: (i) low redshift ($z \sim 0.03 - 0.15$) galaxies containing large amounts of hot dust (53%), including three

❖ Search for anomalous, extrema or “rare” objects in large public datasets

A. Solarz, + 2020, A&A, 642, A103

Planned contributions to GUAIX/IPARCOS activities

❖ Contract María Zambrano (atracción de talento) June 2022-2024



- ❖ Contribute to planning and management of CATARSIS survey with TARSIS (previous talk by A. Gil de Paz)
- ❖ Contribute to software development for processing, post-processing, visualization and analysis of CATARSIS data.
- ❖ Explore synergies with other members of the IPARCOS multidisciplinary team

